

# BepiColombo

ESA cornerstone

## Discovering planet Mercury



bepicolombo

## The Mission

Launch: October 2018

Launcher: Ariane 5  
Launch site: Kourou French Guiana



**MPO**  
(Mercury Planetary Orbiter)

Mass: 1180 kg  
Orbit: 480x1500 km  
Orbital Period: 2h36min  
Data Rate: up to 350 kbps

The **MPO**, built by the European Space Agency, is a masterpiece of technology.

**11 on-board instruments** will work together to collect as much information as possible concerning the interior, the surface, the exosphere, the magnetic and the gravitational field of the planet.

The **MPO** is designed to withstand the extreme environment of Mercury, characterised by a strong solar radiation (up to 13,000 W/m<sup>2</sup>), comparable to about 20 times the radiation received at midday in a desert on the Earth.

The **scientific phase of the mission** will last one Earth year, with a possible extension to a second year.

**Cruise**  
The spacecraft will spend **7 years** on its journey to Mercury, carrying out nine fly-bys: the first one with the Earth, two with Venus and six with Mercury.

**MCS**  
(Mercury Composite Spacecraft)

During the journey to Mercury, **MPO** and **MMO** will be "packed" into the **MCS**, which includes a solar shield and a transfer module (**MTM**) with innovative ion engines.

**MMO - みお**

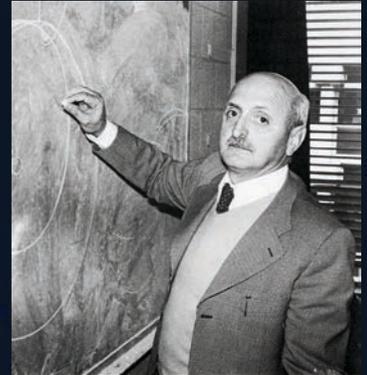
Mass: 288 kg  
Orbit: 890x11640 km  
Orbital Period: about 9h30min

The **Mercury Magnetospheric Orbiter**, built by **JAXA**, the Japanese space agency, embarking 5 instruments, will gather information on the magnetosphere, the exosphere and on the interplanetary medium around Mercury. The **MMO** will spin around itself every 4 seconds, to investigate, in a 360-degree view, the electromagnetic fields, the particles and the interplanetary dust around the planet.

## Giuseppe Colombo

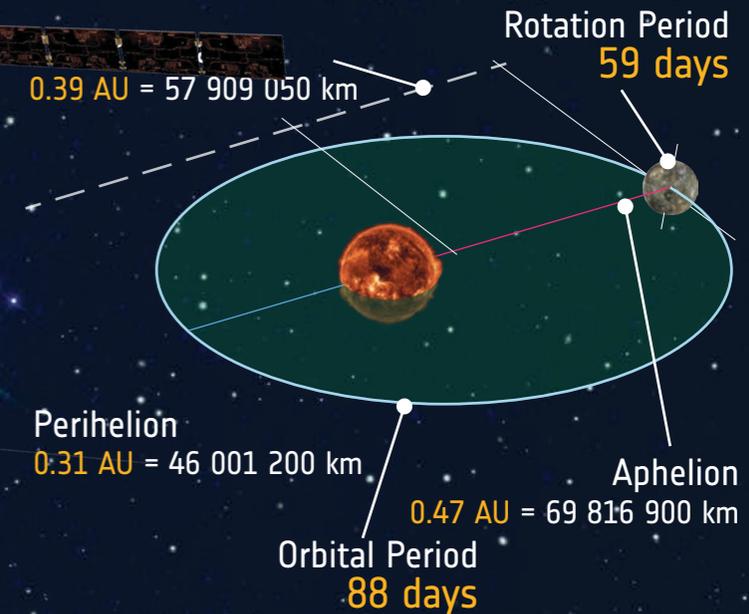
also known as "Bepi" (1920-1984) was a Paduan mathematician and engineer, specialised in celestial mechanics.

He discovered Mercury's spin-orbit 3:2 resonance and contributed to the calculation of the **NASA Mariner 10** probe's orbit to visit Mercury with three fly-bys.



Following its contribution to Mercury science and exploration, **ESA** decided to name the mission designed to discover the mysteries of the Planet.

## Mercury Orbit



Prepared by



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# BepiColombo: discovering planet Mercury

## Relativity

Mercury is the closest planet to the Sun. It is so close that the space-time curvature caused by the mass of the Sun on Mercury's orbit produces measurable effects. The most apparent one is the perihelion precession, which is known since the 19th century. However, the **General Relativity Theory** cannot be considered an ultimate theory, since it is not in accordance with quantum mechanics.

**MORE** will test the **General Relativity Theory** by measuring the space-time curvature with an accuracy of one part per million.

## Our Sun

The radiation of the Sun at Mercury's distance is up to **13,000 W/m<sup>2</sup>**: ten times larger than at the Earth. The flux of photons and particles coming from the Sun is so intense that it can divert the **MPO** trajectory.

**ISA** will measure this kind of "thrust" acting on the **MPO** due to the large flux of photons coming from the Sun. At the same time, the **SERENA** ionised particles detectors will measure the solar wind composition, while the **SIXS** spectrometers will focus on the protons, electrons and solar X-ray fluxes.

## Surface

Widespread lava flows characterise the planet's crust. These volcanic plains are "wrinkled" because of the surface global contraction due to core solidification. The surface can reach **+430°C** when hit by sunlight, and **-170°C** on its darkest and coldest places.

**SIMBIO-SYS** will study the surface in detail to understand the nature of the volcanic plains and their geological evolution with two cameras and a spectrometer.

**BELA**, the laser altimeter, will detect the surface altimetry and reconstruct its topography.

**MERTIS**, an infrared spectrometer, will investigate the surface composition.

**MGNS** will study the radioactive isotopes of the surface and will help finding ice deposits.

Also **MIXS**, an **X-ray** spectrometer, will help uncover the mineralogical composition of the surface.

## Exosphere

The atmosphere of Mercury is so tenuous that its atoms never collide. However, it is so dynamic that can suggest us how Mercury's surface is eroded by the Sun radiation or by micrometeorite impacts.

**SERENA** will provide information on the surface-exosphere-magnetosphere coupled system and the processes involved with 2 units of complementary neutral particle detectors.

**PHEBUS** will investigate the exosphere composition with 2 ultraviolet spectrometers.

## Interior

The planet's core occupies 80% of a total diameter of 4880 km. Its liquid outer part is slowly becoming colder.

**MORE** will investigate the interior structure of Mercury and estimate the size and density of its core by studying the gravitational field. It can also help studying the crust-mantle boundary depth by integrating its data with those of **BELA**.

**ISA** will help **MORE** to account for the effects of **non-gravitational accelerations**, in order to reconstruct Mercury's gravitational field with high accuracy.

## Magnetosphere

Mercury has a dipolar magnetic field, similar to that of the Earth. However the field is too weak to protect Mercury from the hulking presence of the Sun and the strong solar wind.

**MAG** will measure intensity of Mercury magnetic field and, together with **MMO** instruments, will help reconstructing a 3D map of the field.

**SERENA** will provide information on the surface-exosphere-magnetosphere coupled system and the processes involved with 2 units of complementary ionised particle detectors.

**SIXS** will measure the X-radiation and amount of protons and electrons coming from the Sun.

## Four instruments over eleven are headed by Italy



**ISA** - (Italian Spring Accelerometer) is a three-axis accelerometer able to detect the tiny effects ( $10^{-8} \text{ m/s}^2$ ) of the **non-gravitational perturbations** affecting the spacecraft orbit.

The accelerometer allows to take into account all these effects in order to reconstruct the Mercury gravity field.



**MORE** - (Mercury Orbiter Radioscience Experiment) is a Radioscience experiment that, by using on-ground antennas and an on-board Ka-band translator (**KaT**), will allow to measure the Earth-related distance and velocity of the spacecraft with an accuracy of the order of tenths of cm and micron/s, respectively. The goal is to reconstruct the **gravitational field**, to study the **Mercury interior** and to carry out a test of General Relativity.



**SERENA** - (Search for Exospheric Refilling and Emitted Natural Abundances) is a suite of 4 detectors of **neutral atoms** and **ions** which will study the environment around Mercury, by measuring the composition and dynamics of the **exosphere**. In such a way, it will be understood how this tenuous atmosphere is generated and how it interacts with the surface, with the magnetic field of Mercury and with the particles and the solar radiation as well.



**SIMBIO-SYS** - (Spectrometer and Imagers for MPO BepiColombo Integrated Observatory SYStem) is a suite of **three optical instruments**: a **high resolution camera** for a detailed study of Mercury's surface features, a **stereo-camera**, which will make a global 3D reconstruction of the surface and color composite images and a Visible-NIR spectrometer for studying the **surface composition**.

